## The T2K Experiment

Christopher Hearty University of British Columbia/IPP May 21, 2009

## Outline

- Neutrino Oscillations
- T2K Overview
- J-PARC
- Neutrino Beam Line
- Near Detector
- Super-K
- Outlook

## Neutrino Oscillations

- Not expected. One of the biggest scientific discoveries of the last decade.
- Well established.
- Mass eigenstates  $(v_1, v_2, v_3)$  are not the same as weak eigenstates  $(v_e, v_u, v_\tau)$ .

## Simple Two-Generation Case

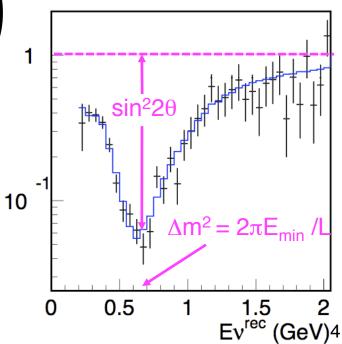
 Probability that a neutrino of energy E with initial flavor α is observed to still be of flavor α after traveling distance L:

$$P(v_{\alpha} \rightarrow v_{\alpha}) = 1 - \sin^{2} 2\theta \cdot \sin^{2} \left( \frac{\Delta m^{2} L}{4E} \right)$$
  

$$\theta \text{ is a mixing angle}$$
  

$$\Delta m^{2} = \Delta m_{b}^{2} - \Delta m_{a}^{2}$$
  

$$can be negative$$



#### **Three Generations**

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{\rm PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$
$$U_{\rm PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & +c_{23} & +s_{23} \\ 0 & -s_{23} & +c_{23} \end{pmatrix} \begin{pmatrix} +c_{13} & 0 & +s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & +c_{13} \end{pmatrix} \begin{pmatrix} +c_{12} & +s_{12} & 0 \\ -s_{12} & +c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$
$$\sim \begin{pmatrix} 0.8 & 0.5 & ??? \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix} \qquad s_{ij} = \sin \theta_{ij}, \ c_{ij} = \cos \theta_{ij}$$

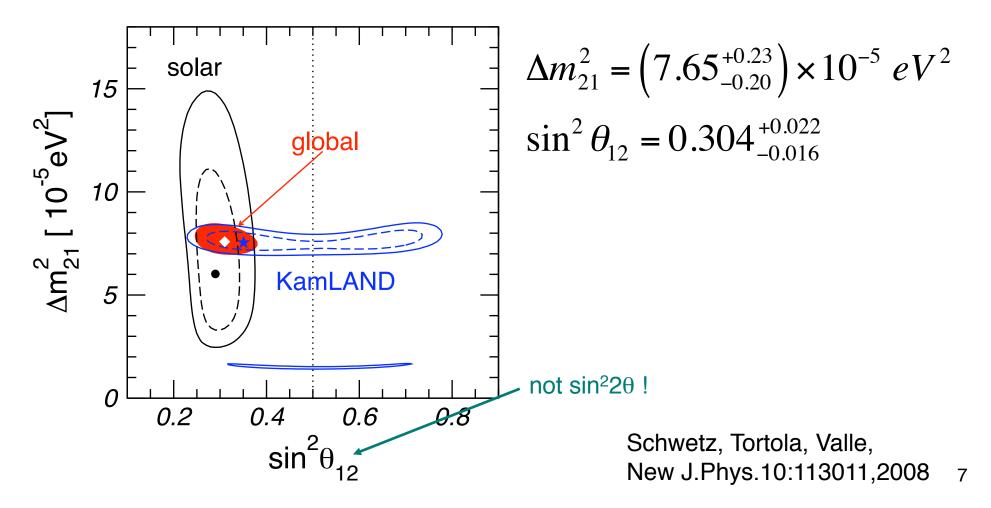
#### This is not a diagonal matrix!

## **Oscillation Parameters**

- 3 mixing angles,  $\theta_{12}$ ,  $\theta_{23}$ ,  $\theta_{13}$
- 1 cp-violating phase  $\delta$
- 3 masses  $m_1$ ,  $m_2$ ,  $m_3 \Rightarrow$  two independent  $\Delta m^2$

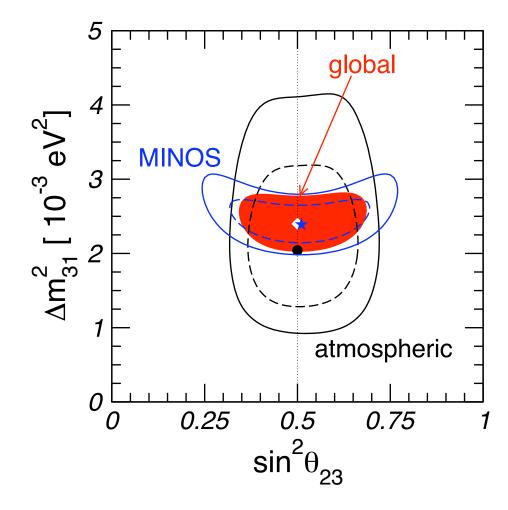
### What We Know – Solar

Super-K, SNO, Gallex, GNO, Kamland (reactor)



## What We Know – Atmospheric

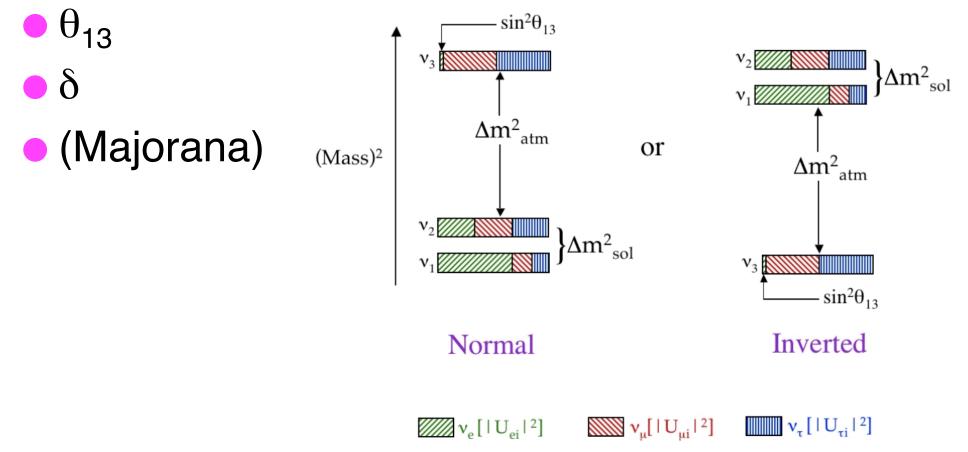
#### Super-K, (K2K), MINOS (FNAL long baseline)



$$\left|\Delta m_{31}^2\right| = \left(2.40_{-0.11}^{+0.12}\right) \times 10^{-3} \ eV^2$$
  
 $\sin^2 \theta_{23} = 0.50_{-0.06}^{+0.07}$ 

## What We Don't Know

• sign of  $\Delta m_{31}^2 \equiv m_3^2 - m_1^2$  (mass hierarchy)



# Measuring $\theta_{13}$ via $v_{\mu} \rightarrow v_{e}$ Oscillations

- Dominant "atmospheric" oscillation is  $v_{\mu}$  to  $v_{\tau}$ .
  - »  $\tau$  are not reconstructed, so this is basically  $\nu_{\mu}$  disappearance.
- ν<sub>µ</sub> → ν<sub>e</sub> can occur, but has not been observed.
   » complicated formula...

$$P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$$
where

$$T_1 = \sin^2 \theta_{23} rac{\sin^2[(1-x)\Delta]}{(1-x)^2}$$
 ,

$$T_2 = \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} ,$$
  
$$T_3 = \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} ,$$

$$T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$$

$$\begin{split} &\Delta \equiv \Delta m_{31}^2 L / 4E \\ &\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2 \approx \pm 0.03 \\ &x \equiv 2\sqrt{2}G_F N_e E / \Delta m_{31}^2 \approx \pm E / 11 GeV \\ &N_e = \text{electron number density} \end{split}$$

$$P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$$

where

$$T_1 = \sin^2 \theta_{23} \frac{\sin^2[(1-x)\Delta]}{(1-x)^2} , \qquad T_1 \text{ is dominant, } \sim 0.5$$
  
at oscillation max

$$T_2 = \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)}$$
 ,

$$T_3 = \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} ,$$

$$T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$$

and

$$\begin{split} &\Delta \equiv \Delta m_{31}^2 L/4E \end{split} \qquad \text{kinematic phase of oscillation} \\ &\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2 \approx \pm 0.03 \end{aligned} \qquad \alpha \text{ is small} \\ &x \equiv 2\sqrt{2}G_F N_e E / \Delta m_{31}^2 \approx \pm E / 11 GeV \\ &N_e = \text{electron number density} \end{split}$$

 $P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$ where

$$T_1 = \sin^2 \theta_{23} rac{\sin^2[(1-x)\Delta]}{(1-x)^2}$$
 ,

$$T_2 = \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)}$$

$$T_3 = \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)}$$

 $T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$  .

This term does not depend on 
$$\theta_{13}$$
, but is small

,

,

and

$$\begin{split} &\Delta \equiv \Delta m_{31}^2 L \big/ 4E \\ &\alpha \equiv \Delta m_{21}^2 \big/ \Delta m_{31}^2 \approx \pm 0.03 \\ &x \equiv 2\sqrt{2}G_F N_e E \big/ \Delta m_{31}^2 \approx \pm E \big/ 11 GeV \\ &N_e = \text{electron number density} \end{split}$$

$$P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$$
  
where

$$T_1 = \sin^2 \theta_{23} \frac{\sin^2[(1-x)\Delta]}{(1-x)^2}$$
 ,

$$\begin{split} T_2 &= \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \quad , \\ T_3 &= \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \quad , \end{split}$$

$$T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$$

$$\begin{split} &\Delta \equiv \Delta m_{31}^2 L/4E \\ &\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2 \approx \pm 0.03 \\ &x \equiv 2\sqrt{2}G_F N_e E / \Delta m_{31}^2 \approx \pm E/11 GeV \end{split} \qquad \text{Matter effects are larger for} \\ &N_e = \text{electron number density} \qquad (E \sim 0.6 \text{ GeV}) \end{split}$$

$$P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$$
where

$$T_1 = \sin^2 \theta_{23} \frac{\sin^2[(1-x)\Delta]}{(1-x)^2} \quad ,$$

$$T_2 = \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} ,$$
  
$$T_3 = \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} ,$$

$$T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$$

$$\begin{split} &\Delta \equiv \Delta m_{31}^2 L/4E \\ &\alpha \equiv \Delta m_{21}^2/\Delta m_{31}^2 \approx \pm 0.03 \\ &x \equiv 2\sqrt{2}G_F N_e E/\Delta m_{31}^2 \approx \pm E/11 GeV \\ &N_e = \text{electron number density} \end{split}$$

All three of these terms depend on the mass hierarchy (sign of  $\Delta m_{31}^2$ )

$$P(\nu_{\mu} \rightarrow \nu_{e}) \cong \sin^{2} 2\theta_{13} T_{1} - \alpha \sin 2\theta_{13} T_{2} + \alpha \sin 2\theta_{13} T_{3} + \alpha^{2} T_{4}$$
where
$$T_{1} = \sin^{2} \theta_{23} \frac{\sin^{2}[(1-x)\Delta]}{(1-x)^{2}} \quad \delta \text{ only appears in product with } \sin 2\theta_{13}$$

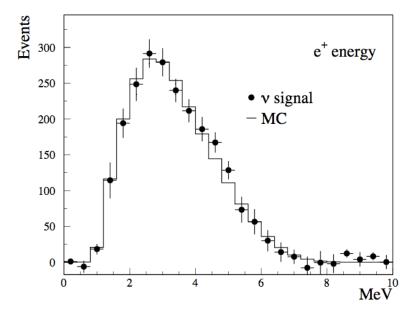
$$T_{2} = \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \qquad CP \text{ violating } \delta \rightarrow -\delta \text{ for } \overline{\nu}$$

$$T_{3} = \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \qquad CP \text{ conserving } \delta \rightarrow -\delta \text{ for } \overline{\nu}$$

$$T_4 = \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2}$$

# Chooz Experiment

- Reactor experiment (1997-98) looking for ν<sub>e</sub> disappearance. E<sub>ν</sub> ~ 3 MeV, L ~ 1 km.
   P(v
  <sub>e</sub> → v
  <sub>e</sub>) = 1 sin<sup>2</sup>(2θ<sub>13</sub>) · sin<sup>2</sup>(Δm<sup>2</sup><sub>31</sub>L/4E) + O(α<sup>2</sup>)
  - » Does not depend on  $\delta$ , *x*, or mass hierarchy.

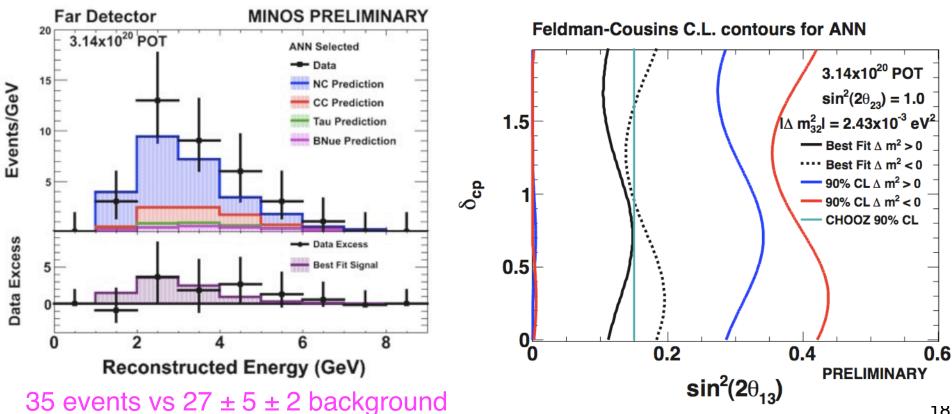


 $\sin^2 2\theta_{13} < 0.15$  90% CL

Double Chooz (upgrade with two detectors) aiming for sensitivity to ~0.03

## MINOS Results on $\theta_{13}$

• v beam from NuMI (120 GeV),  $E_v \sim 3$  GeV. Far detector in Soudan mine, 735 km. 4 kT fiducial.



## Physics Goals of T2K

- Measure  $\theta_{13}$
- Improve measurements of  $\theta_{23}$  and  $\text{Im}_{31}^2$ I
- So is this the equivalent of a dedicated experiment to measure V<sub>ub</sub>?

## Grand Unified Theories

 $V_{\rm PMNS} = \begin{pmatrix} 0.837 e^{-i179^{\circ}} & 0.544 e^{-i173^{\circ}} & 0.0566 e^{i138^{\circ}} \\ 0.364 e^{-i3.86^{\circ}} & 0.609 e^{-i173^{\circ}} & 0.705 e^{i3.45^{\circ}} \\ 0.408 e^{i180^{\circ}} & 0.577 & 0.707 \end{pmatrix}$ **MNS Matrix:**  $\sin^2 heta_{
m atm} = 1, \ \tan^2 heta_{\odot} = 0.422 \ \ {
m and} \ |U_{e3}| = 0.0566$ 

prediction for Dirac CP phase:  $\delta = -46.9$  degrees  $J_{\ell} = -0.0094$ 

Note that these predictions do NOT depend on u and  $\xi_0$ 

neutrino masses:

 $u_0 = -0.0593, \quad \xi_0 = 0.0369, \quad M_X = 10^{14} \text{ GeV}$  $m_1 = 0.0156 \text{ eV}, \quad m_2 = 0.0179 \text{ eV}, \quad m_3 = 0.0514 \text{ eV}$ 

2 parameters in neutrino sector

 $\alpha_{21} = \pi$ Majorana phases

$$\alpha_{31} = 0.$$

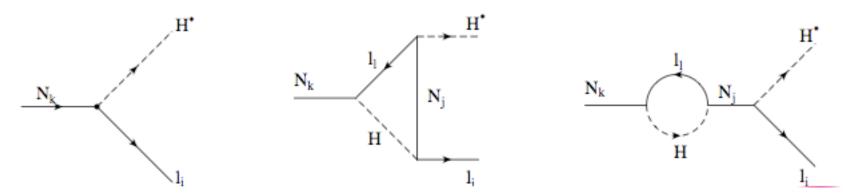
predicting:3 masses, 3 mixing angles, 3 CP Phases

41

Mu-Chun Chen, Phys. Lett. B652 (2007) 34 20

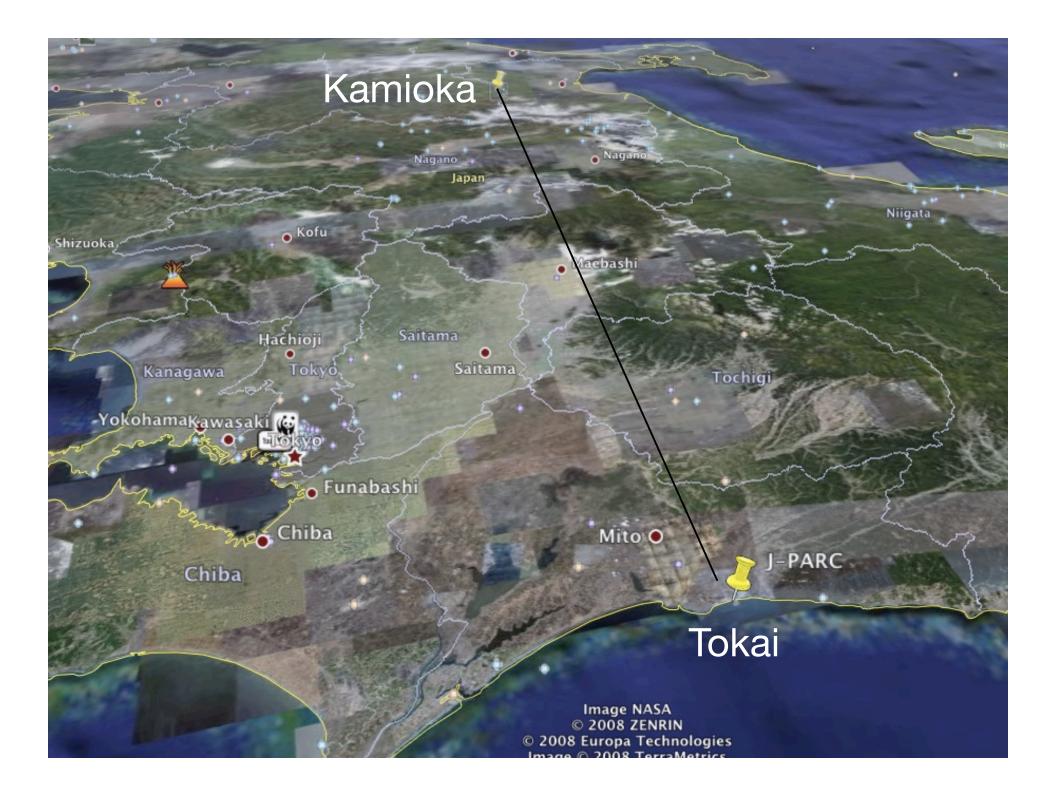
# Leptogenesis

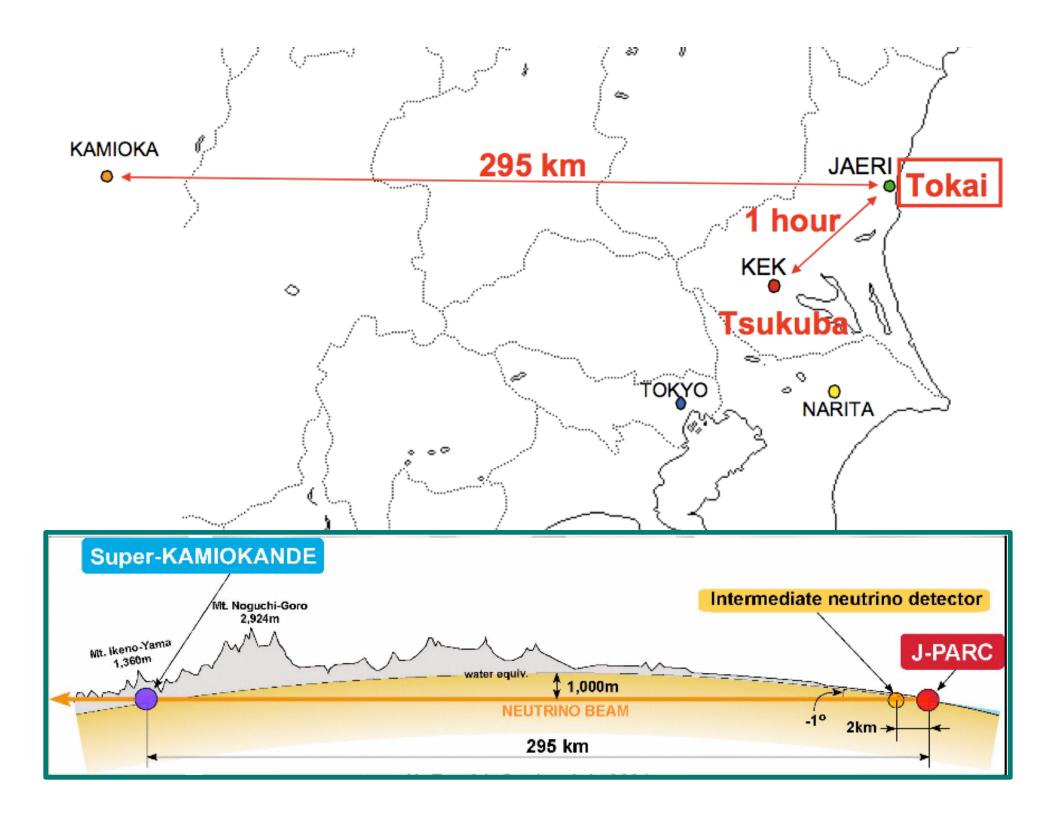
- A CPV measurement could shed light on matter anti-matter asymmetry of the universe.
  - Standard Leptogenesis: decays of RH neutrinos (CPV in decay) Quantum interference of tree diagram and one-loop diagram



 There are models in which strong correlations exist between low-energy and high-energy mixing parameters.
 M.-C.Chen & K.T. Mahanthappa, Phys. Rev. D71 (2005) 035001

#### **T2K OVERVIEW**





#### **T2K Collaboration**

Antarctica

#### • ~400 people, (290 PhD physicists)

• Japan (85)

 ICRR, Hiroshima U, KEK, Kobe U, Kyoto U, Miyagi U of Education, Osaka City U, U of Tokyo

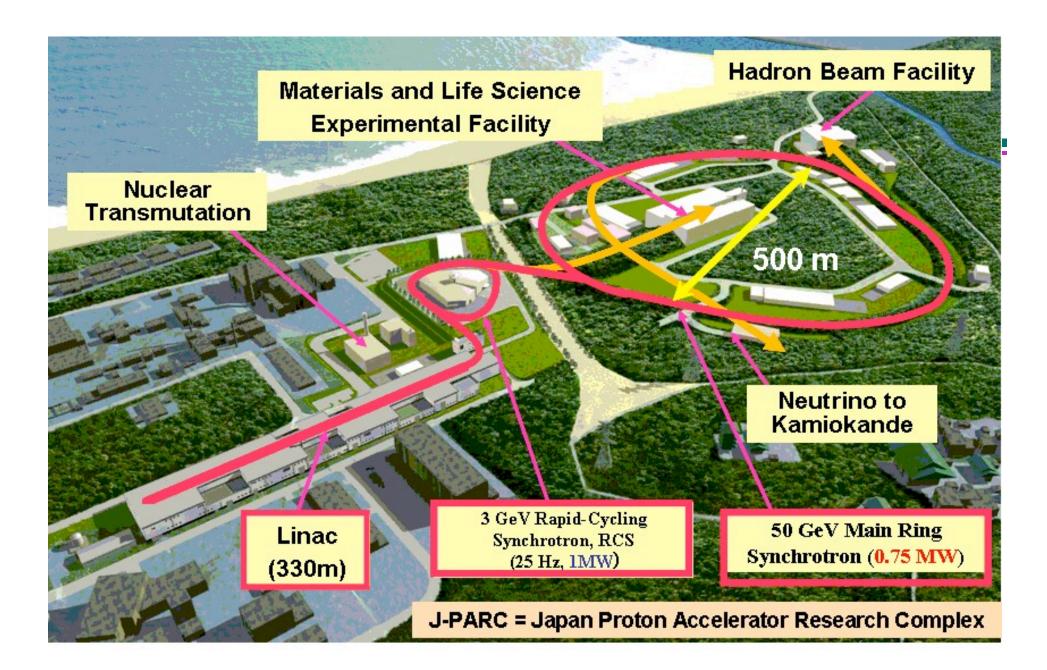
- UK (83)
  - Oxford, Imperial College London, Lancaster U, Queen Mary U of London, Sheffield U, STFC/RAL/Daresbury Lab, U of Liverpool, U of Warwick
- U.S.A. (66)
  - Boston U, Brookhaven Lab, Colorado State U, Duke U, Louisiana State U, Stony Brook U, UC Irvine, U of Colorado, U of Pittsburgh, U of Rochester, U of Washington
  - Canada (65)
    - U of British Columbia, U of Regina, TRIUMF, U of Toronto, U of Victoria, York U
  - France (51)
    - CEA/DAPNIA Saclay, IPN Lyon, LLR Ecole Polytechnique, LPNHE-Paris
  - Switzerland (38)
    - Bern, ETHZ, U of Geneva

- Poland(29)
  - IFJ PAN Cracow, IPJ Warsaw, Technical University
     Warsaw, U of Silesia, Warsaw U, Wroclaw U
- Russia (13)
  - **INR**
- Spain(II)
   IFIC Valencia, Barcelona/IFAE
- Italy (10)
  - ▶ INFN-Bari, INFN-Rome, Napoli, Padova, Rome
- Korea (9)
  - Chonnam National U, Dongshin U, Sejong U, Seou
     National U, Sungkyunkwan U
- Germany(3)
  - RWTH Aachen U

## T2K Overview

- Measure  $\theta_{13}$  using  $v_{\mu} \rightarrow v_{e}$  oscillations. » also significantly lower errors on  $\sin^{2}2\theta_{23}, \Delta m_{23}^{2}$
- Experiment has three components:
  - » Neutrino beam line, using 30 GeV p from J-PARC
  - » Near detector ND 280
     » Far detector Super-K
     Not the same technology!
- $1^{st}$  stage: 100 kW ×  $10^7$  sec [2 ×  $10^{20}$  p.o.t.]  $2^{nd}$  stage: 1 MW ×  $10^7$  sec [2 ×  $10^{21}$  p.o.t.] full data set: 3.5 MW ×  $10^7$  sec [7 ×  $10^{21}$  p.o.t.]

### **THE J-PARC FACILITY**



Joint KEK / JAEA project



#### Tracks

- 1. Godzilla Attacks Tokai Japan's Energy Crisis (M2)
- 2. Fateful Confrontation (M3)
- 3. Main Title (M4)



#### Accelerator Chain

#### Linac: Accelerates H<sup>-</sup> to 180 MeV



#### Accelerator Chain

- Rapid-cycling synchrotron RCS: 1 MW of 3-GeV protons (mostly for spallation neutrons)
  - » also serves as injector for main ring.

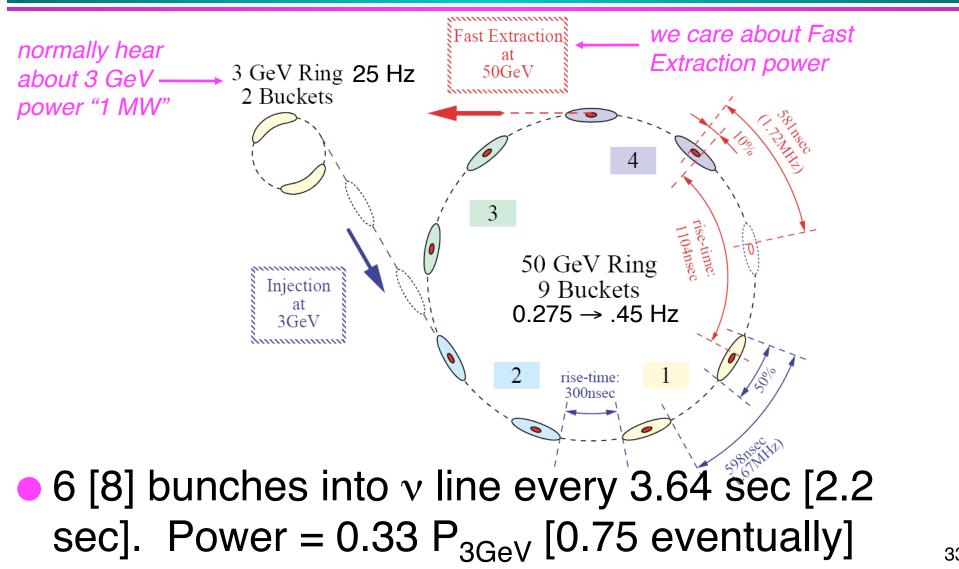


#### Accelerator Chain

 Main Ring: synchroton accelerates from 3 GeV to 30 GeV. Protons for v beamline (fast extraction) or for Kaon physics (slow extraction)



### Note on Power

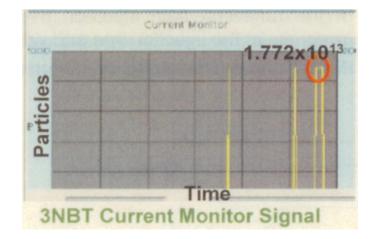


#### Status



### Status

- 3 GeV RCS: 210 kW for 70 sec
  - » limited by beam dump license.



- Single bunch operation of MR at 30 GeV; fast extraction to v line.
- However, front end of linac (Radio Frequency Quadrupole) is not so healthy...

# **RFQ Status**

- Accelerates H<sup>-</sup> to 3 MeV.
- Discharges started in Sept 08. Perhaps related to incorrect copper alloy, and/or inadequate vacuum (protons).



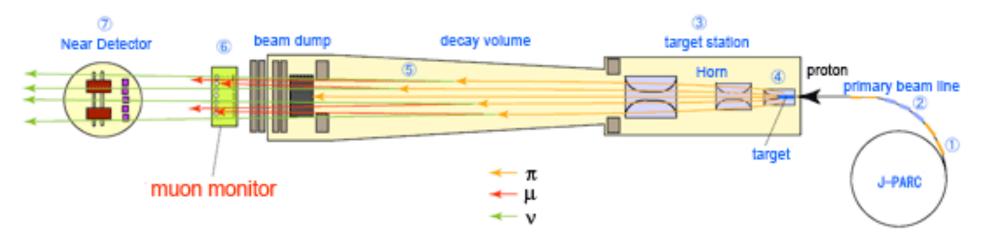


## RFQ Status II

- Currently limiting RCS (3 GeV) power to 20 kW
- Conditioning 5 hours per day; stable operation for 19 hours.
- Extra pumping being installed
- Replacement RFQ aimed for Spring 2010.

### **NEUTRINO BEAM LINE**

## Target Area



- 30 GeV protons strike 90 cm graphite target, creating π<sup>±</sup> and K<sup>±</sup>.
- Three horns focus  $\pi^+$ , defocus  $\pi^-$
- $\pi^+ \rightarrow \mu^+ \nu_{\mu}$  in 110 m decay volume
- Muon monitor follows beam dump.



#### 2<sup>nd</sup> Horn

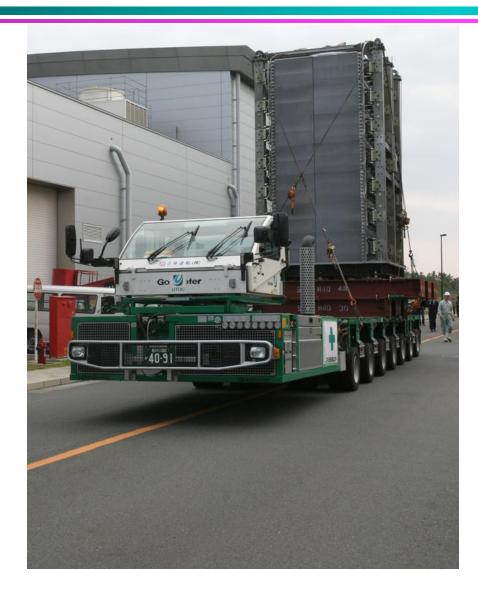


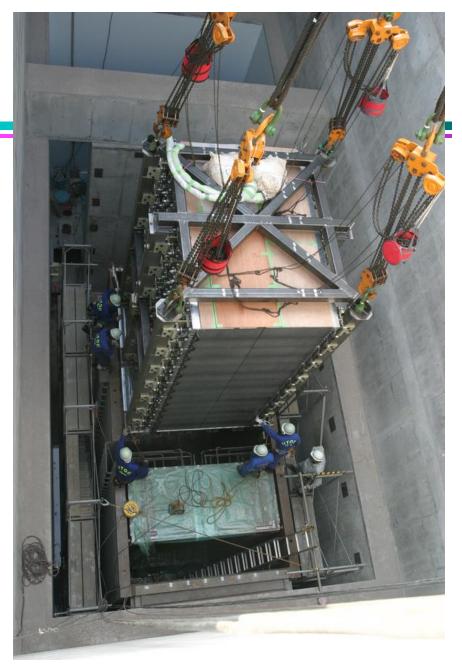
# Inside the decay volume





#### Beam Dump



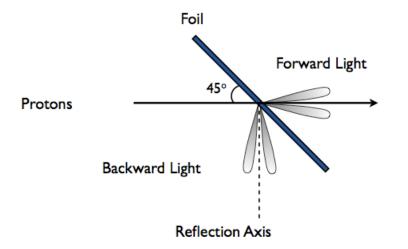




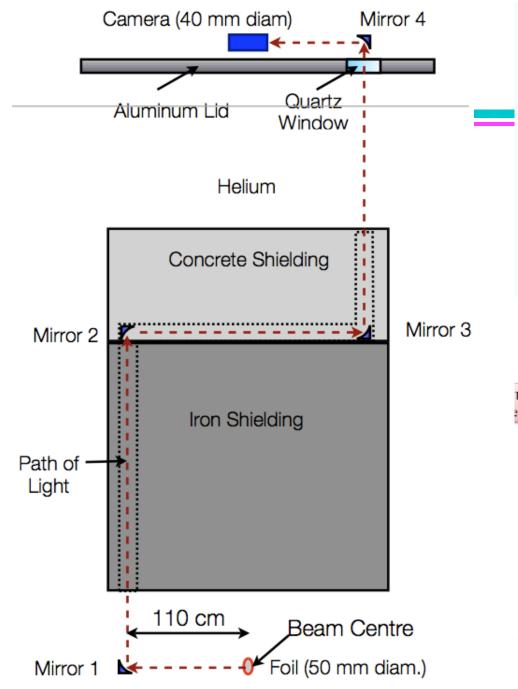
Muon Monitor 10<sup>8</sup> μ/cm<sup>2</sup>/spill at full power (after beam dump)

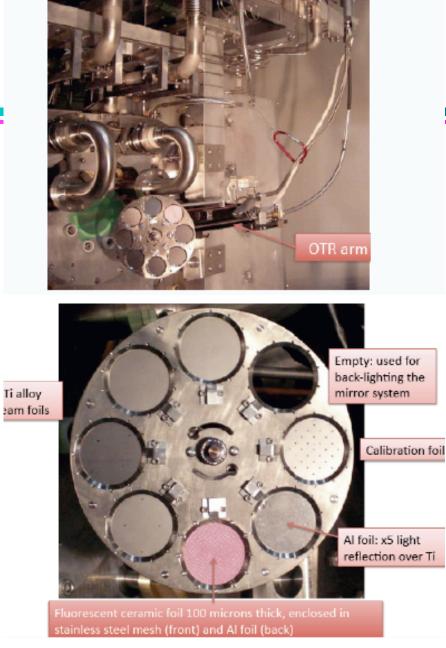
## Monitoring the Beam Location

- Horns focus point-to-direction. Need to monitor beam location to ~1 mm to ensure correct v beam direction.
- Optical Transition Radiation detector immediately in front of target.



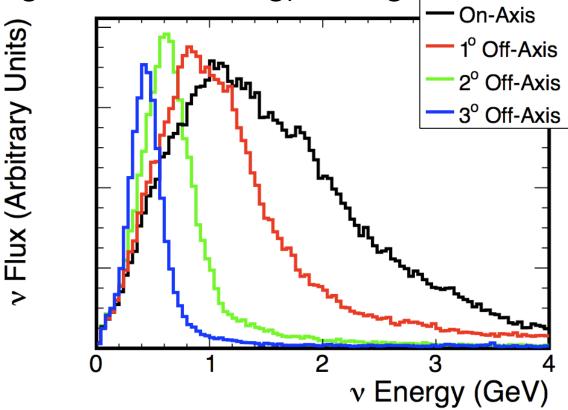
#### Cross Section Looking Downstream





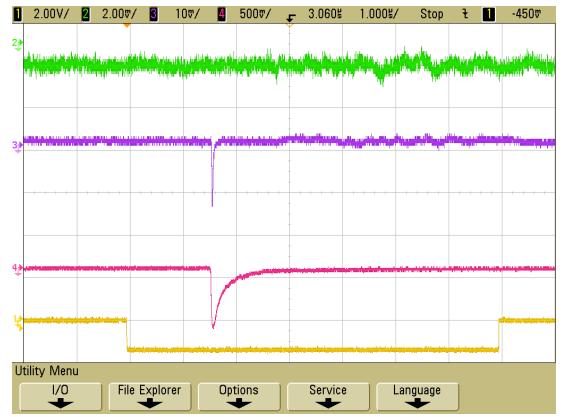
### Off Axis Beam

- Super-K is 2.5° off the v beam direction
  - » higher flux at energy of interest; lower flux at higher (background-inducing) energies.



#### Status

#### First beam on target April 23, 2009



signal in muon monitor

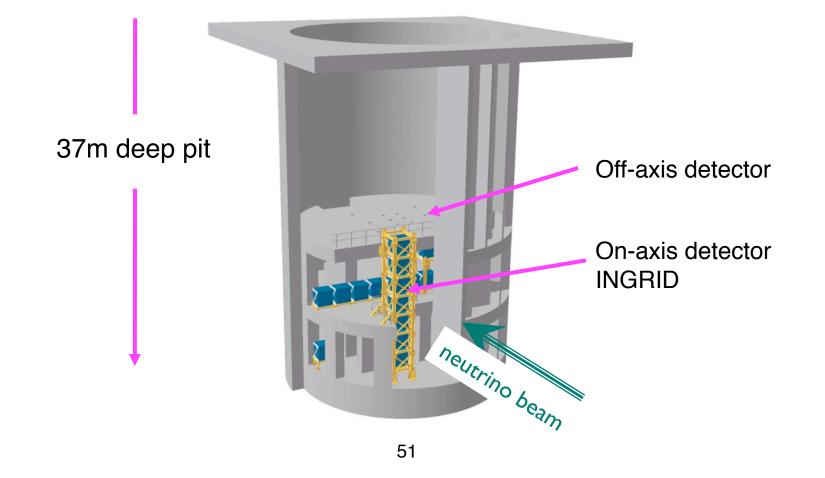
## Commissioning

- Commission with low intensity, single pulses until official government inspection on May 28.
   » activation
- 2<sup>nd</sup> and 3<sup>rd</sup> horns then installed over next four months. Last remaining major components.
- Commission at higher power in the fall.

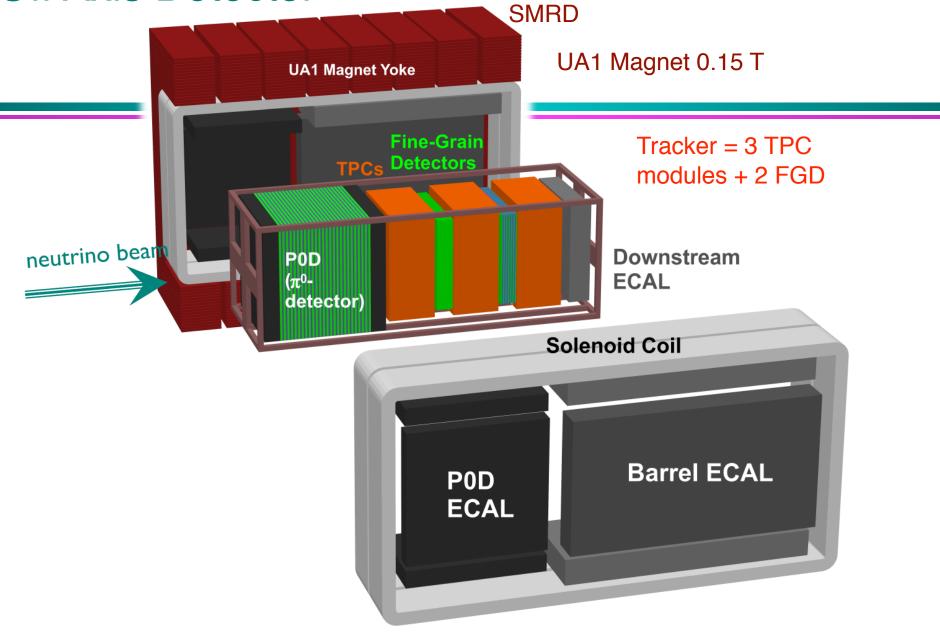
### **NEAR DETECTOR**

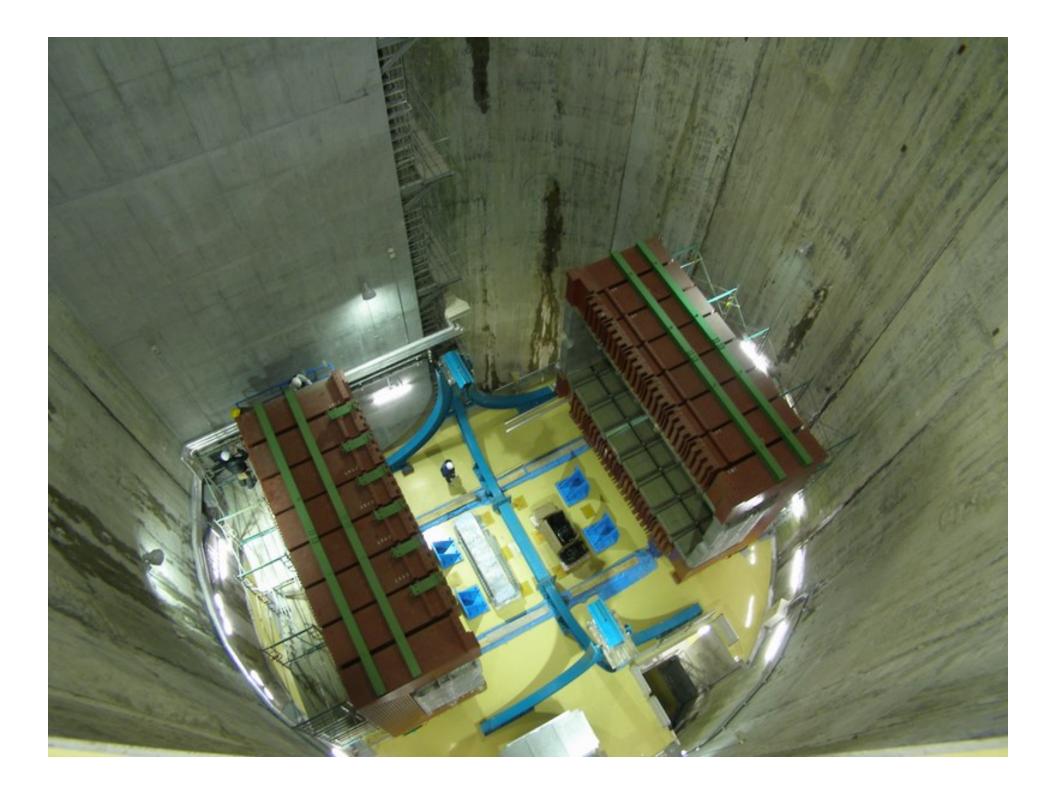
### ND 280

#### On- and off-axis detectors 280 m from target



#### **Off Axis Detector**





### Goals of the Near Detector

- Measure v beam direction
- Determine  $v_{\mu}$  flux and energy spectrum
- Study backgrounds to  $v_e$  appearance ( $\theta_{13}$ )
- Study backgrounds to  $v_{\mu}$  disappearance

## Ingrid – On-Axis Detector

 16 modules × 10 tons; 10<sup>5</sup> interactions/day at full power ⇒ ~1 mrad precision on beam direction in a day.

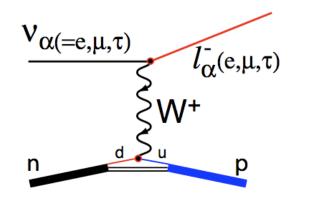


scintillator bars/WLS fiber /iron layers/ Hamamatsu MPPC readout

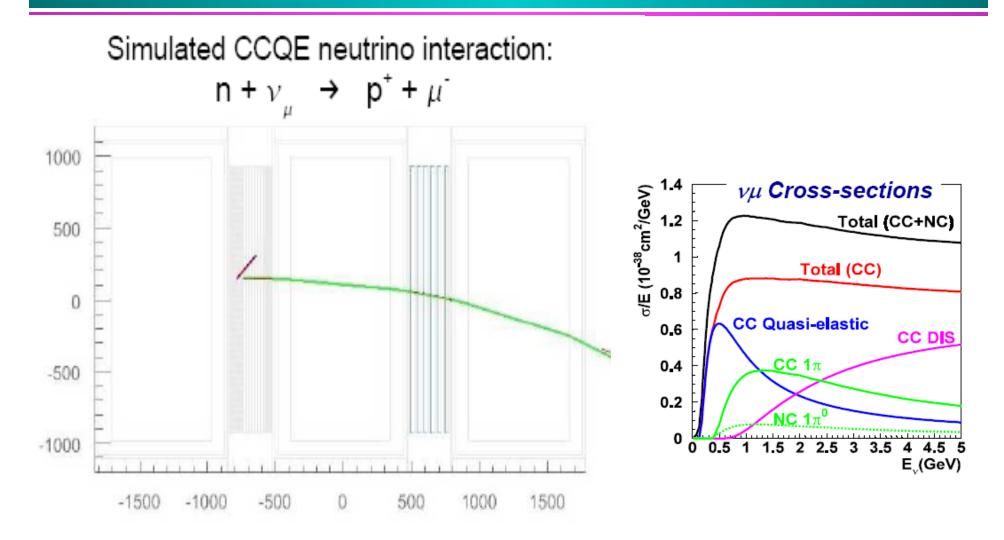


### Quasi-elastic v Interactions

• "Golden Mode" for flux measurements: charged-current quasi-elastic CCQE  $vn \rightarrow p\ell^-$ 



- lepton  $\theta$  + momentum  $\Rightarrow$  neutrino energy, assuming target is at rest.
  - » Not quite true Fermi motion.



### Flux Measurements

- ~30k CCQE events in Tracker (FGD + TPC) fiducial volume in initial data set.
- Main background: charged current events with extra  $\pi^+$  (often through  $\Delta$  resonance)
  - » skewed kinematics
  - » constrain with proton angle
- v<sub>e</sub> ~0.5% of beam at ~0.6 GeV (oscillation maximum)

Backgrounds to Measurement of  $\theta_{13}$ Using  $v_e$  Appearance

- Intrinsic  $v_e$  component of beam
- Neutral current production of  $\pi^0$

$$\gg v_{\mu} N \rightarrow v_{\mu} \pi^0 N$$

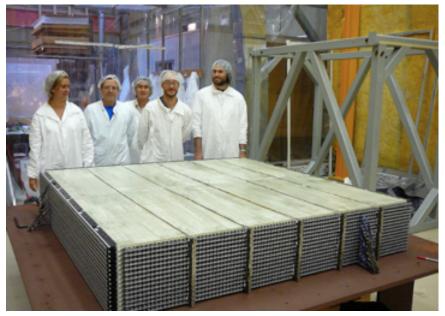
- » one photon is missed, other looks like an electron.
- » typically from neutrinos with energies above oscillation maximum.

### Nuclear Effects

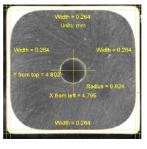
- Non-negligible hadronic interactions in nucleus.
   » rescattering, absorption, charge exchange
  - »  $\pi^{-}$  mostly result from such interactions.
- Neutrino cross sections are not well known at these energies, so measure flux × cross section on water.

### Fine Grain Detector – FGD

- Two modules, each ~1 ton.
  - » 1 cm extruded scintillator bars/WLS/MPPCs
  - » 2<sup>nd</sup> module also has water layers.
  - > small bars  $\Rightarrow$  good proton reconstruction.



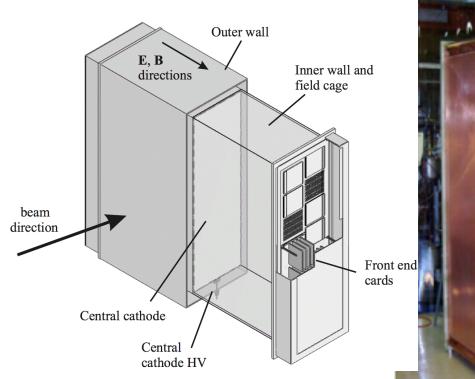




## TPCs

- Three rectangular TPCs with micromegas readout.
  - » δp/p ~10% at 1 GeV is adequate (Fermi motion)

  - » dE/dx gives >3 $\sigma$  µ/e separation 0.3–1 GeV





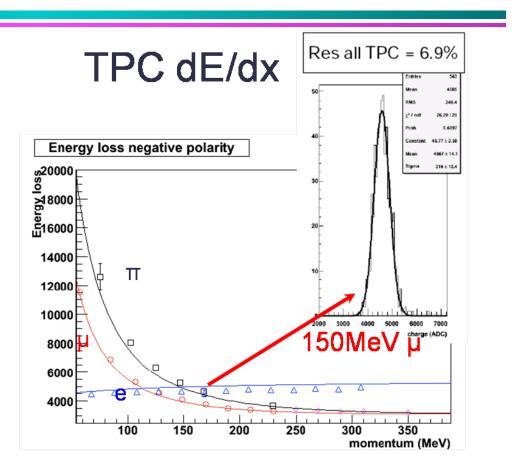




### Tracker Beam Test at TRIUMF



stopping track with decay electron 400 200 TPC FGD -200 in-time hits -400 delayed hits -600 -800 00 -200 200 400 600 800 1000

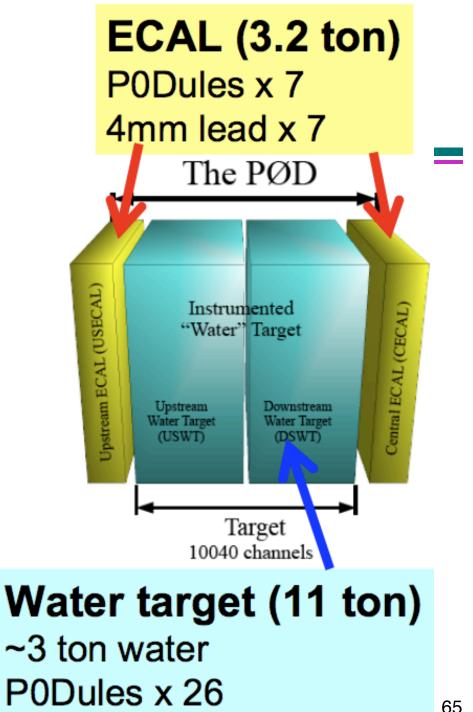


 $>5\sigma$  separation of e/µ for p>200MeV

Performance as expected

### P())

- Optimized to measure neutral current  $\pi^0$ production on water.
- Triangular scintillator bars/WLS fibers/lead/ water





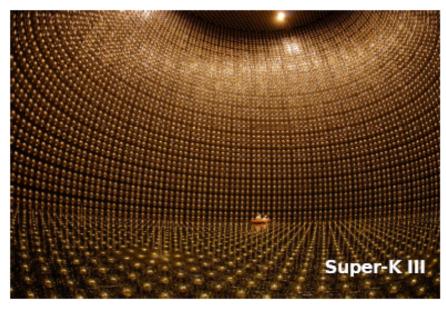
### ND 280 Status

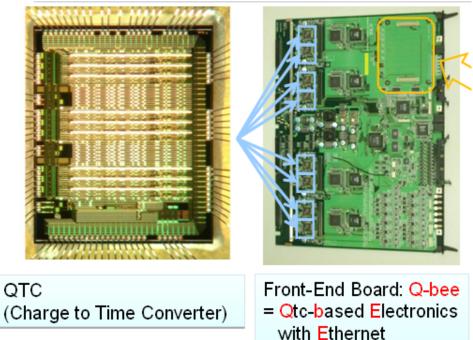
- Infrastructure installation in progress
- Magnetic field mapping in July
- P0D, FGD, downstream ECAL and two TPCs will arrive by July and be installed early October.
  - » 3<sup>rd</sup> TPC should arrive December.
  - » Full side (barrel) ECAL in Summer 2010.

### **FAR DETECTOR**



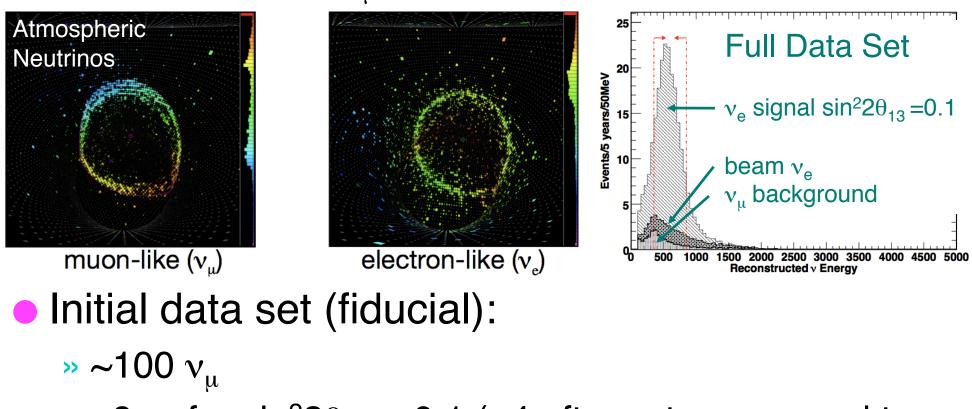
- Large water Cherenkov detector: 50 kT total, 22.5 kT fiducial.
- In operation since 1996; SK-IV started Sept 08 with new electronics.





## Super-K Signals

#### • CCQE for both $v_{\mu}$ and $v_{e}$ spectra



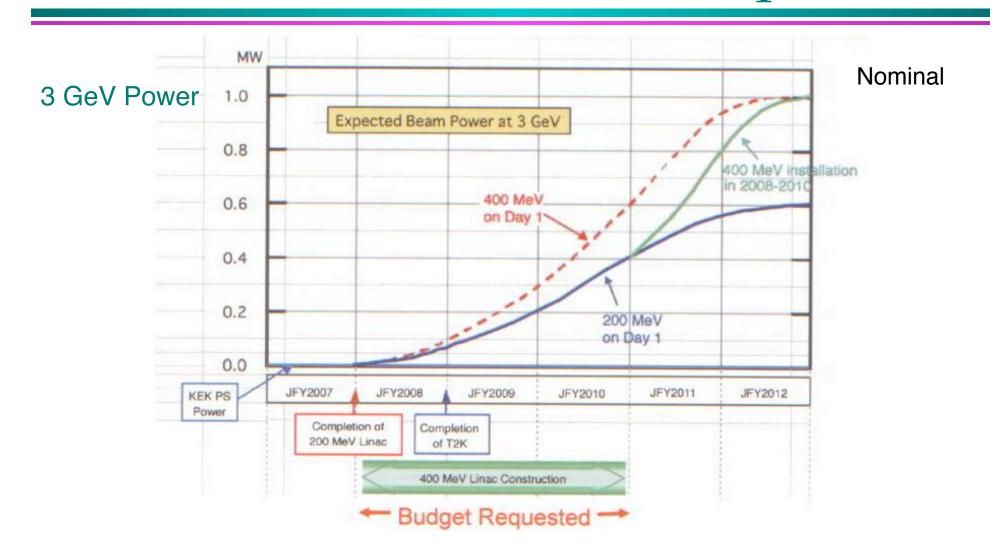
» ~9 v<sub>e</sub> for sin<sup>2</sup>2θ<sub>13</sub> = 0.1 (~4 after cuts compared to ~0.8 background)

### OUTLOOK

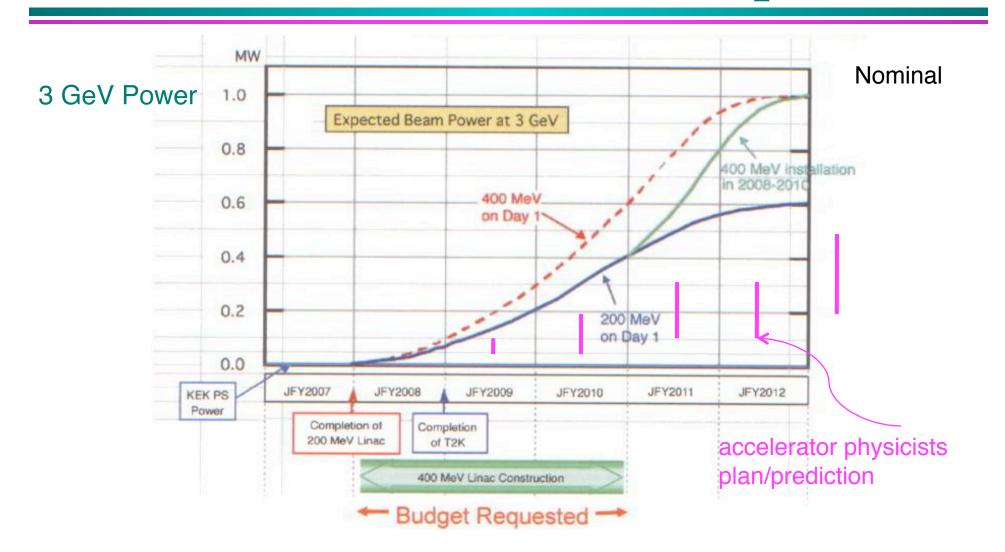
## **Expected Power**

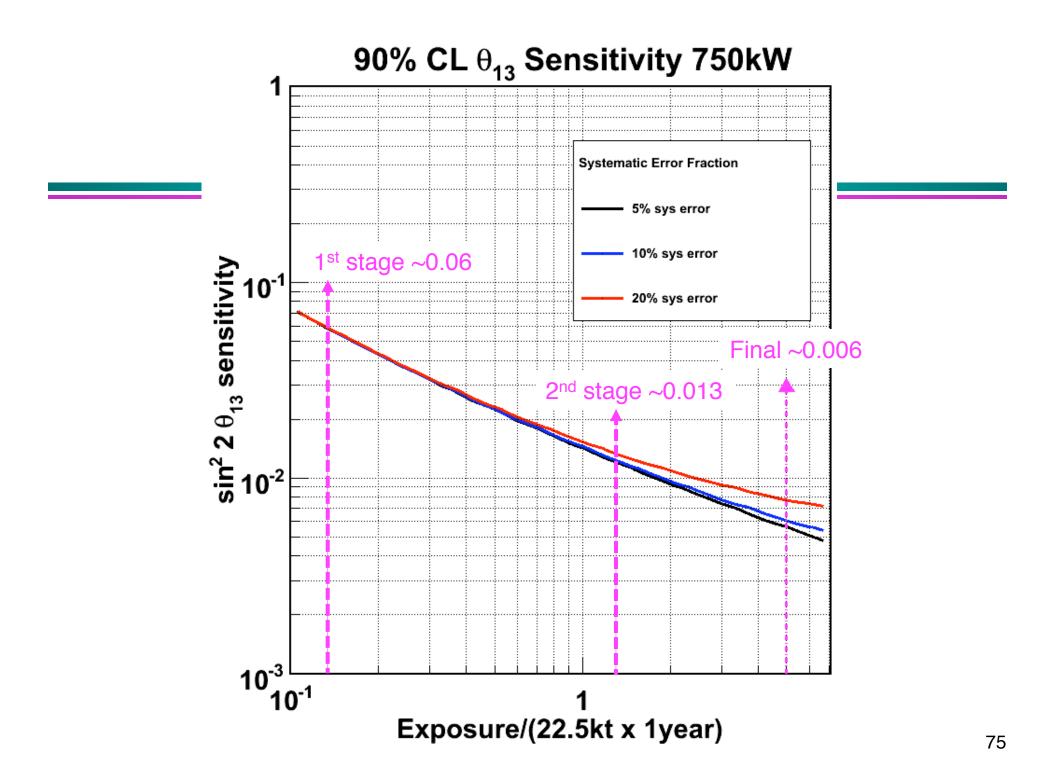
- First run starts Dec/Jan. Not clear if we will get the planned 100 kW on target until RFQ is replaced in Spring 2010.
  - » also split run time between slow and fast extraction.
- It will take a while to get full 750 kW on target (= 1 MW @ 3 GeV).

## Plan for Power Ramp



## Plan for Power Ramp

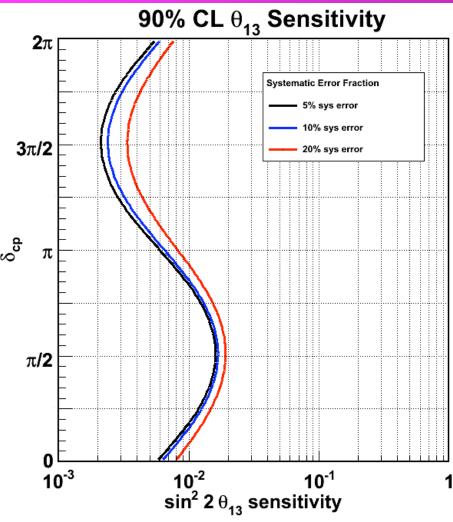




## Dependence on $\delta_{CP}$

• We actually measure a function of  $\theta_{13}$ ,  $\delta_{CP}$  and matter effects/mass hierarchy.

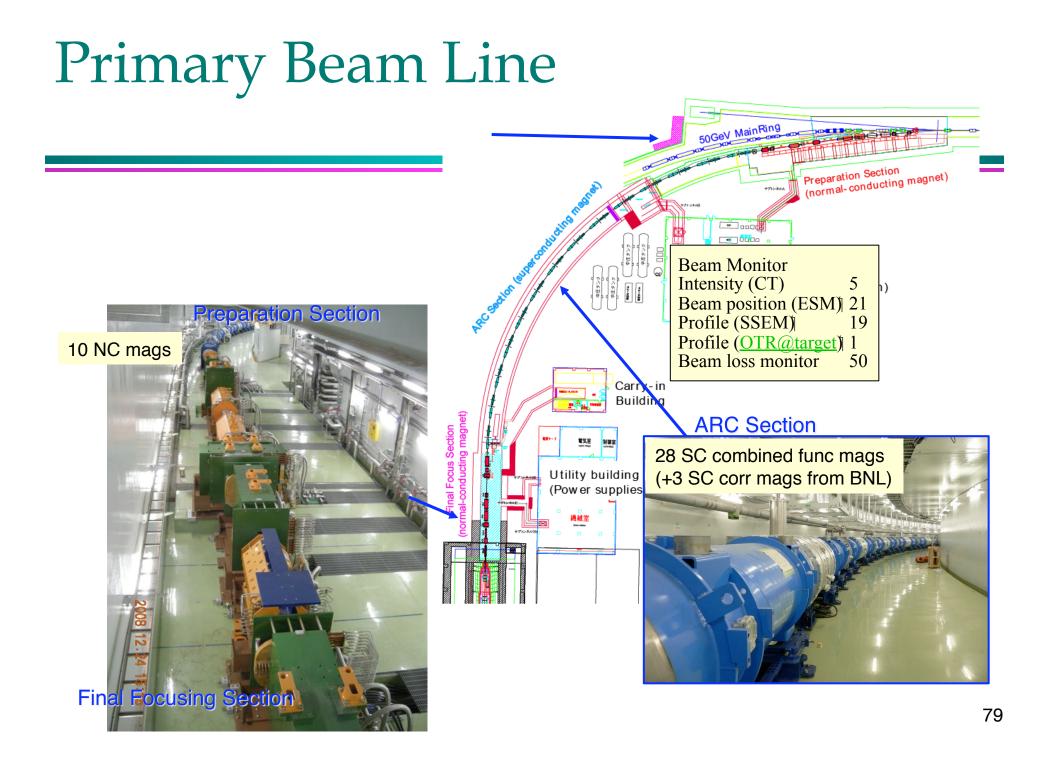
If we are lucky, compare<sup>δ</sup> with Double Chooz (no δ, no x) and NOvA (larger matter effects) to disentangle

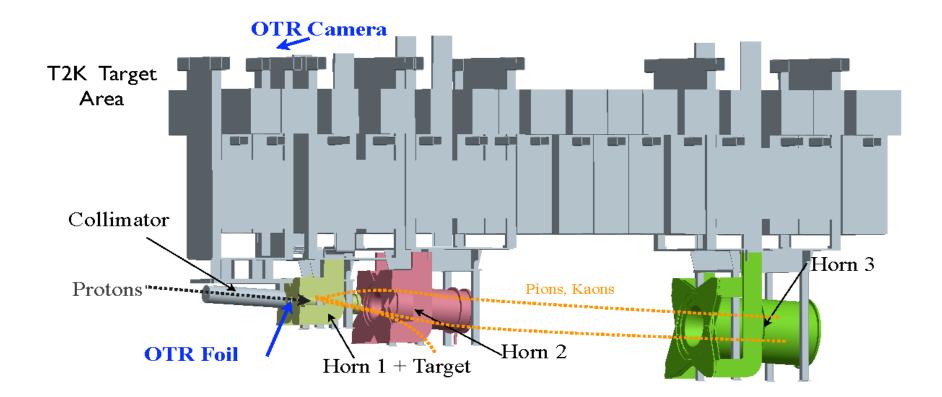


# Summary

- J-PARC and T2K beam line commissioning is in progress and near detector installation will occur this summer.
- 1<sup>st</sup> stage result on  $\theta_{13}$  (Summer 2010? 2011?) should surpass existing limits.
- Full dataset sensitivity to  $\sin^2 2\theta_{13} \sim 0.01$ .
  - » but I hope we are not still talking about a limit at that time!

#### BACKUP





## Tri-bimaximal PMNS Matrix

neutrino masses: triplet Higgs contribution

$$3_S = \frac{1}{3} \begin{pmatrix} 2\alpha_1\beta_1 - \alpha_2\beta_3 - \alpha_3\beta_2\\ 2\alpha_3\beta_3 - \alpha_1\beta_2 - \alpha_2\beta_1\\ 2\alpha_2\beta_2 - \alpha_1\beta_3 - \alpha_3\beta_1 \end{pmatrix} \qquad 1 = \alpha_1\beta_1 + \alpha_2\beta_3 + \alpha_3\beta_2$$

- neutrino masses: singlet contribution  $1 = \alpha_1\beta_1 + \alpha_2\beta_3 + \alpha_3\beta_2$
- Resulting mass matrix:

$$M_
u = rac{\lambda v^2}{M_x} egin{pmatrix} 2\xi_0 + u & -\xi_0 & -\xi_0 \ -\xi_0 & 2\xi_0 & u - \xi_0 \ -\xi_0 & u - \xi_0 & 2\xi_0 \end{pmatrix}$$

$$V_{
u}^{ ext{T}} M_{
u} V_{
u} = ext{diag}(u + 3\xi_0, \, u, \, -u + 3\xi_0) rac{v_u^2}{M_x}$$

$$U_{\text{TBM}} = \begin{pmatrix} \sqrt{2/3} & 1/\sqrt{3} & 0\\ -\sqrt{1/6} & 1/\sqrt{3} & -1/\sqrt{2}\\ -\sqrt{1/6} & 1/\sqrt{3} & 1/\sqrt{2} \end{pmatrix}$$

Form diagonalizable:

- -- no adjustable parameters
- -- neutrino mixing from CG coefficients!

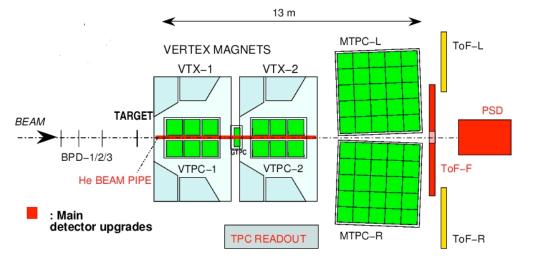
General conditions for Form Diagonalizablility in seesaw: M.-C. C, S. F. King, arXiv:0903.0125

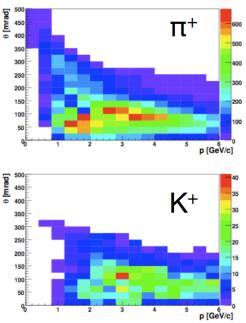
Mu-Chun Chen

## NA-61 at CERN SPS

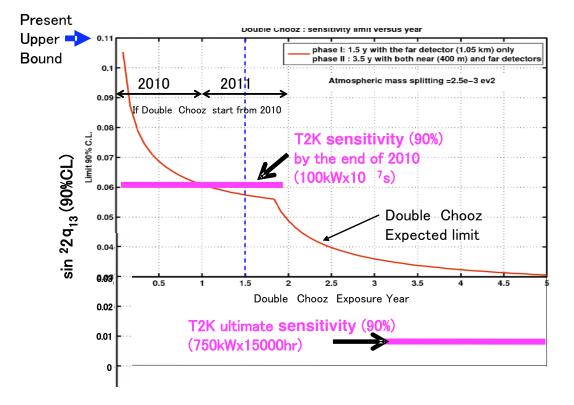
- Beamline group is responsible for calculating the produced v flux and spectrum.
- Results from NA-61 are important input. Hadron production using 30 GeV p on T2K target.

» 3 week run in Aug 2009

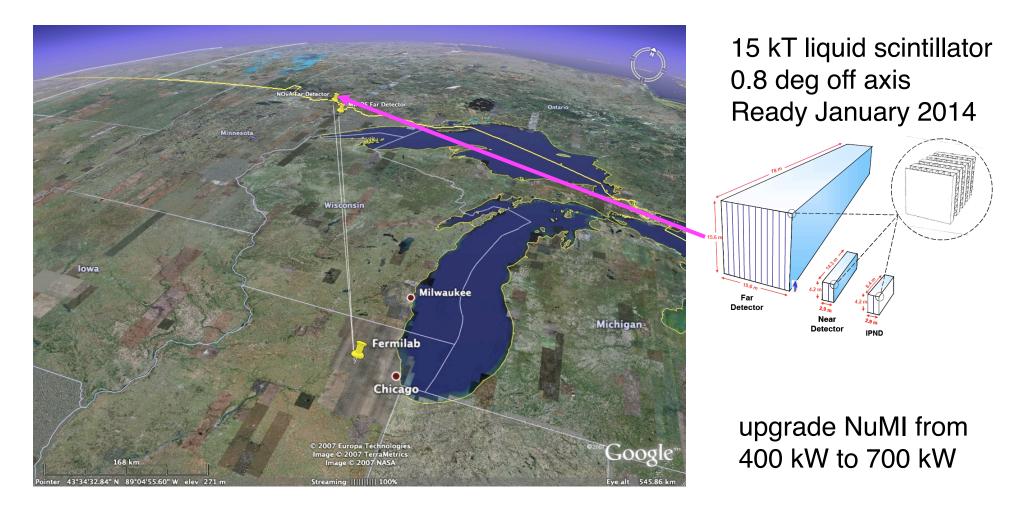




## **Competition from Double Chooz**

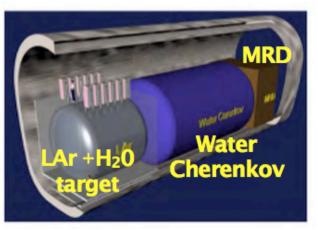


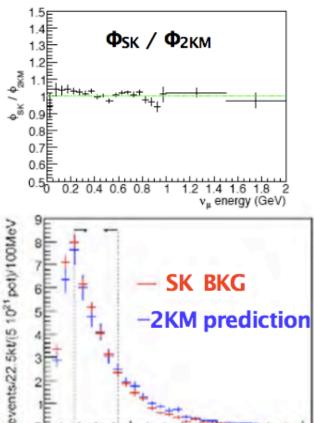
### NOvA



#### 2KM Proposal

- Strengthen prediction of far detector flux through measurement (~2 km away) using :
  - Almost same beam flux as far detector
  - Same target material
  - Same detector technology and reconstruction analysis
- Check ND280+NA61 prediction before oscillation
- Combine ND280+2kM measurements
  - Reduce further the systematic errors
  - Understand better the ve backgrounds





reconstructed v\_ energy (MeV)

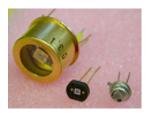
2000

1000

3000

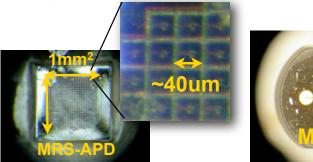
Antonin Vacheret, Imperial College London

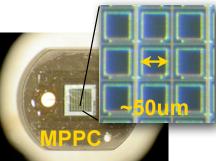
## Novel type of photosensor : Geiger Mode multi- pixels APD



#### ND280 Scintillator Detectors Constraints :

- Magnetic Field
- Very tight space constraints
- Low light yield at end of WLS fibre
- High number of channels
- Detector operation 5 years



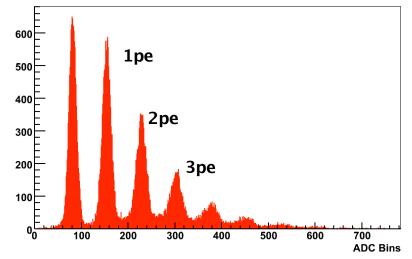


 $Signal = \sum Q_i^{pixel}$ 

- Insensitive to magnetic field, tested to 4T
- Small (active area ~1mm<sup>2</sup>)
- Bias voltage 50V
- Photon detection efficiency >20% at 500nm
- Gain G~ 5x10<sup>5</sup>
- Low power consumption
- Dark count rate ~0.5MHz/mm<sup>2</sup> at 25°C
- Longevity OK

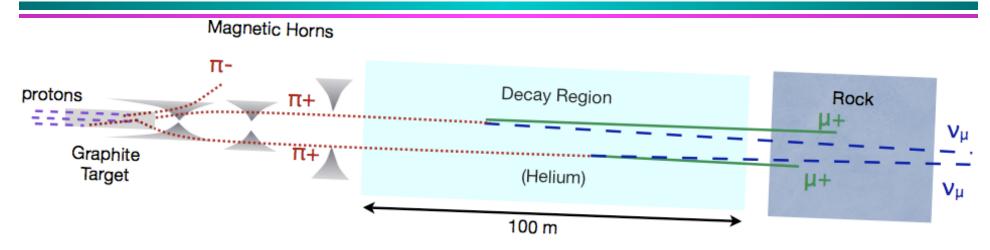
Antonin Vacheret, Imperial College London

Photo-electron resolution + dynamic range combined in 1 sensor



Nulnt07, May 30, June 3 2007 Fermilab

# Target Area



- 30 GeV protons strike 90 cm graphite target, creating π<sup>±</sup> and K<sup>±</sup>.
- Three horns focus  $\pi^+$ , defocus  $\pi^-$
- $\pi^+ \rightarrow \mu^+ \nu_{\mu}$  in 110 m decay volume
- Muon monitor follows beam dump.

# **Reactor experiments (Q13)**

Site (proposal)	Power (GW)	Baseline Near/Far (m)	Detector Near/Far (t)	Overburden Near/Far (MWE)	Sensitivity to sin²2θ <sub>13</sub> (90%CL)	Operation year
Angra dos Reis (Brazil)	6.0	300/1500	50/500	200/1700	~0.006	unknown
Double Chooz (France)	8.7	400/1050	10/10	115/300	~ 0.03	Early 2010
Daya Bay (China)	17.4	360/500/ 1800	40/40/80	264/302/946	~ 0.008	Late 2010
Reno (S. Korea)	17.3	290/1380	15/15	200/540	~ 0.025	Early 2010